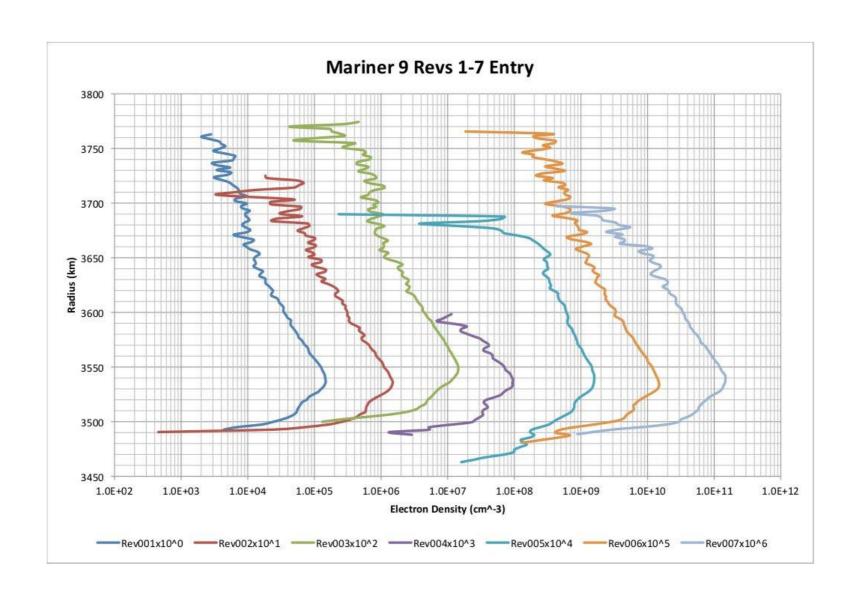
Ionosphere of Mars: Miscellaneous stuff

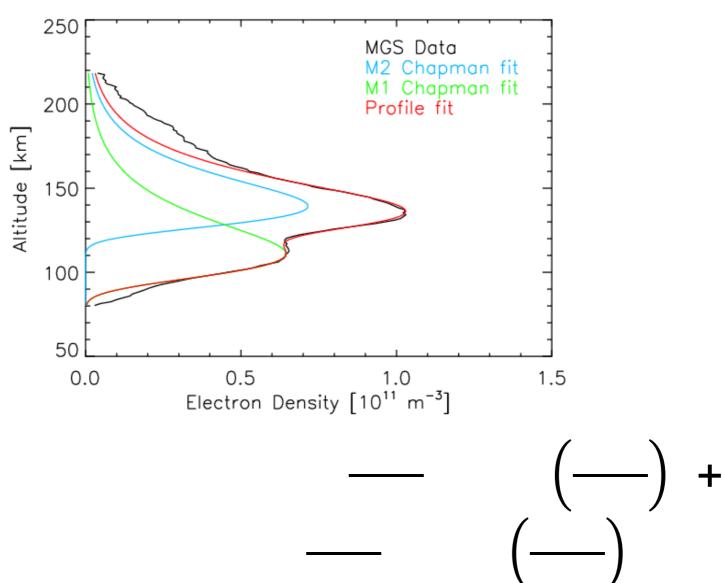
Paul Withers withers@bu.edu

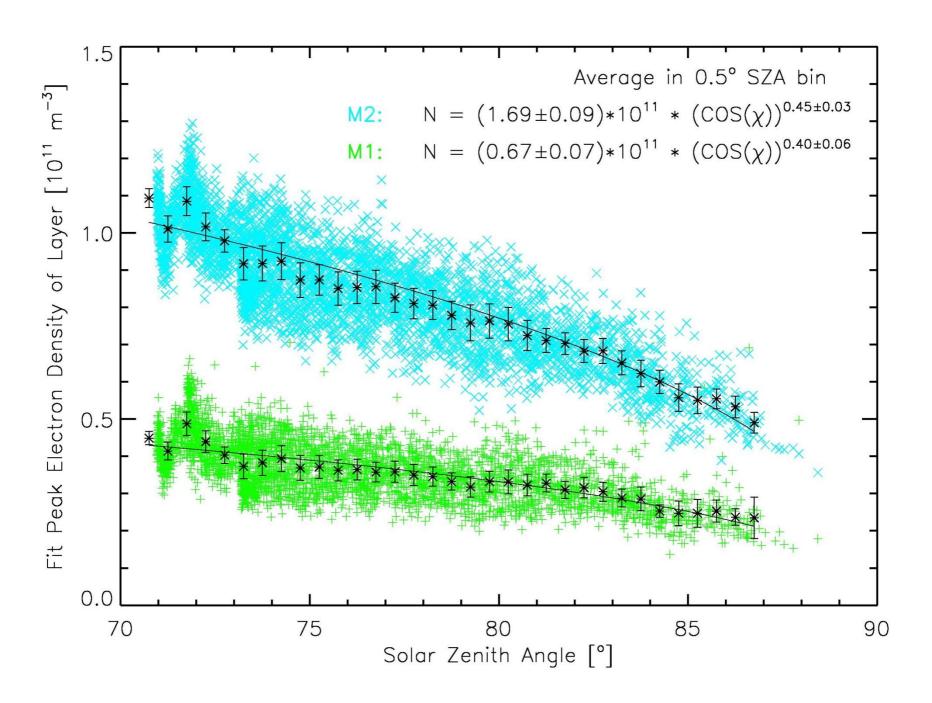
MEX/VEX RS team meeting Cologne, Germany 2012.01.11-13

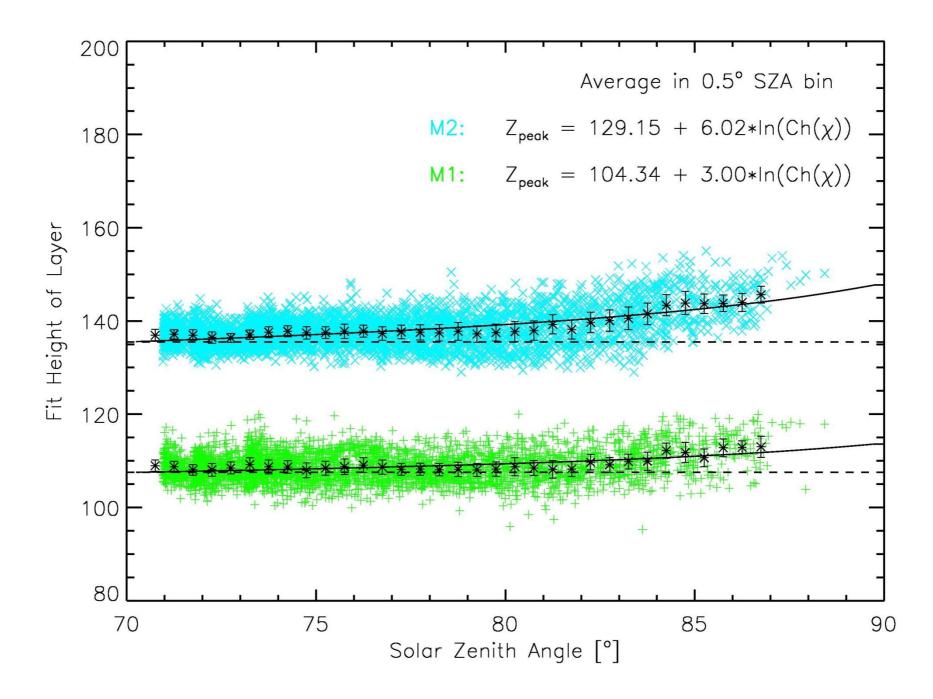
Mariner 9 profiles recovered!

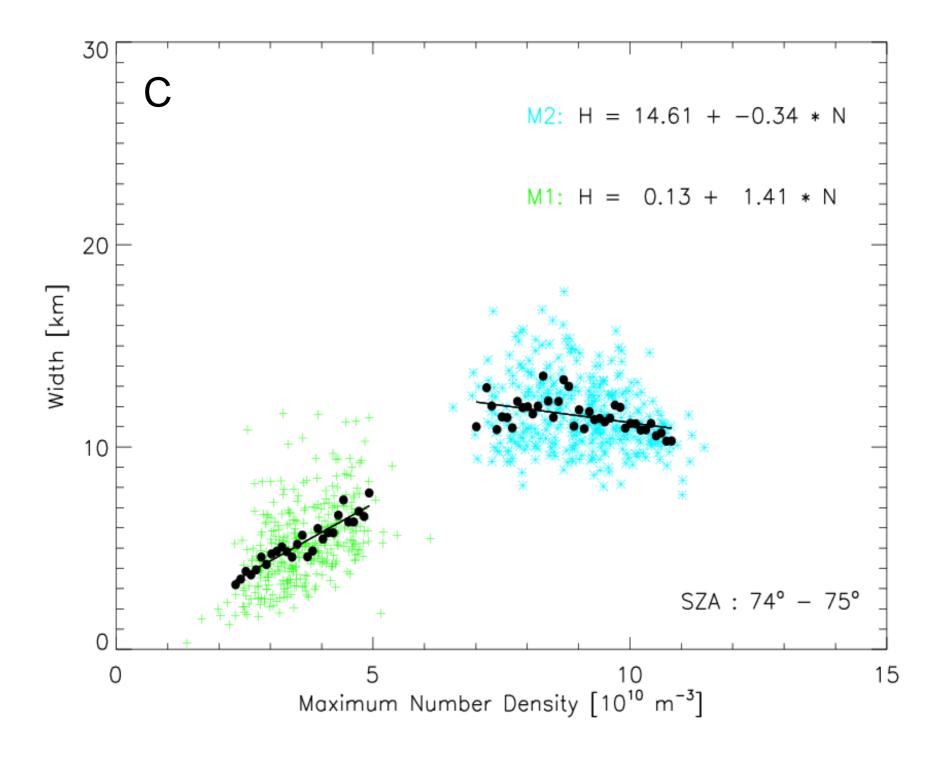


Comparing M1 and M2 layers (Katy Fallows)

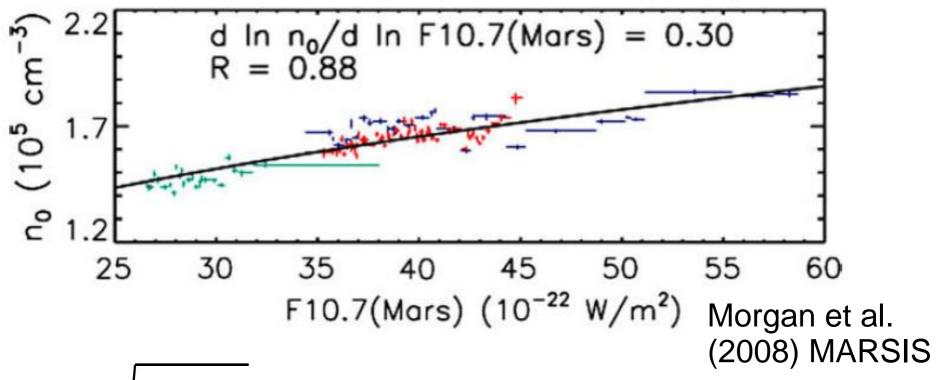






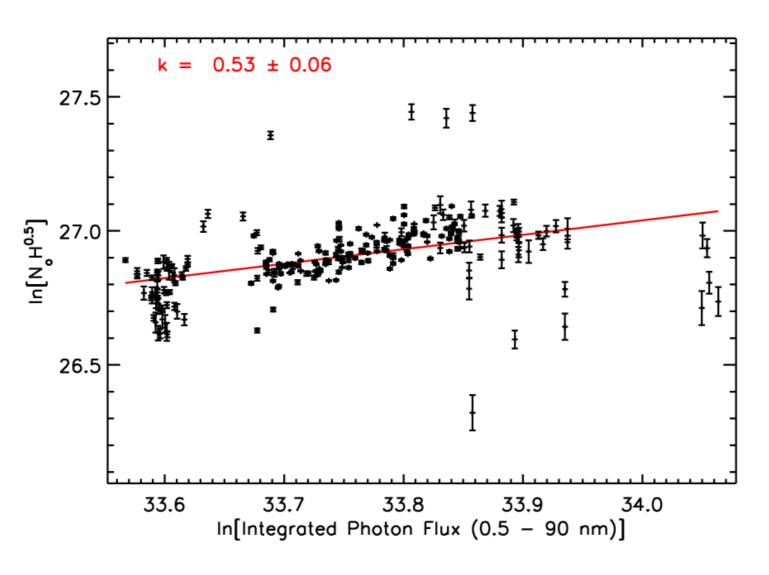


How Chapmanlike? (Zachary Girazian)



$$N_o = \sqrt{\frac{F}{\alpha e H}}$$
 $N_o H^{0.5} \propto F^k$ k should be 0.5

What is best proxy for F?

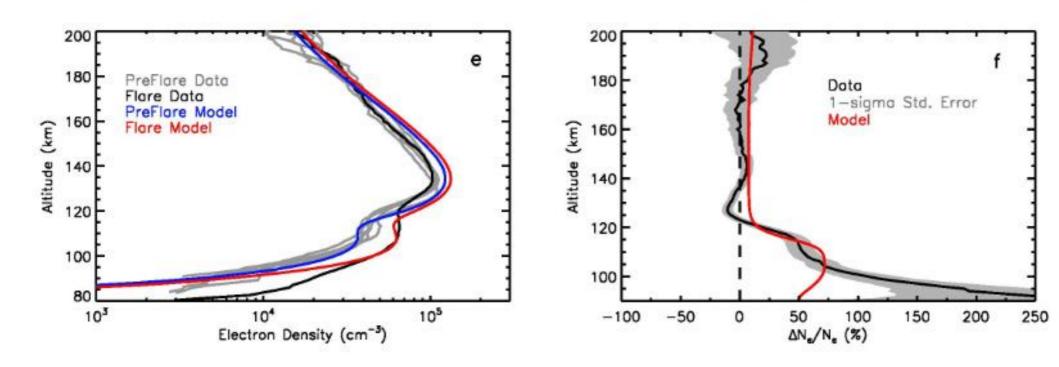


MEX data for SZA < 80

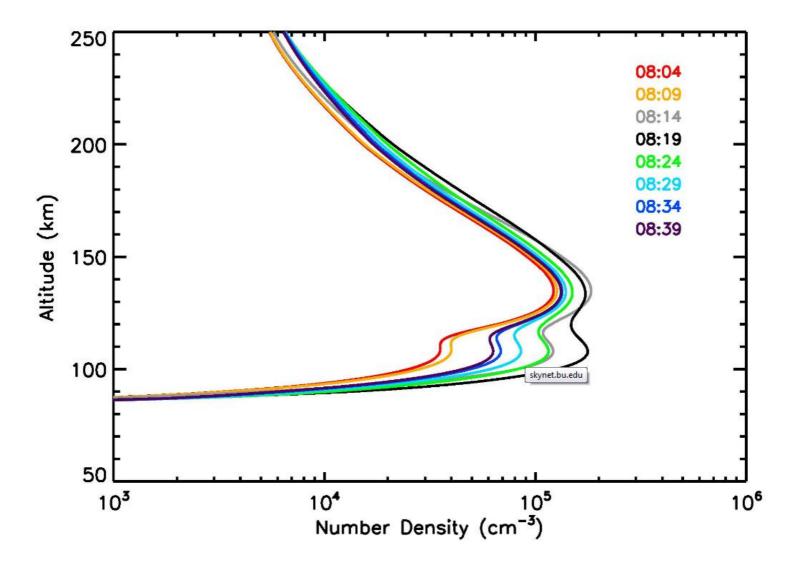
Caution!
Best fit value for k keeps changing whenever we modify anything

Preliminary at best

Response to solar flares



Lollo et al. manuscript submitted to JGR-Space Physics on 22 November



Chemical composition (Majd Matta)

0₂⁺ 0⁺

CO2+

NO[†]

HCO[†]

CO[†]

Ar⁺

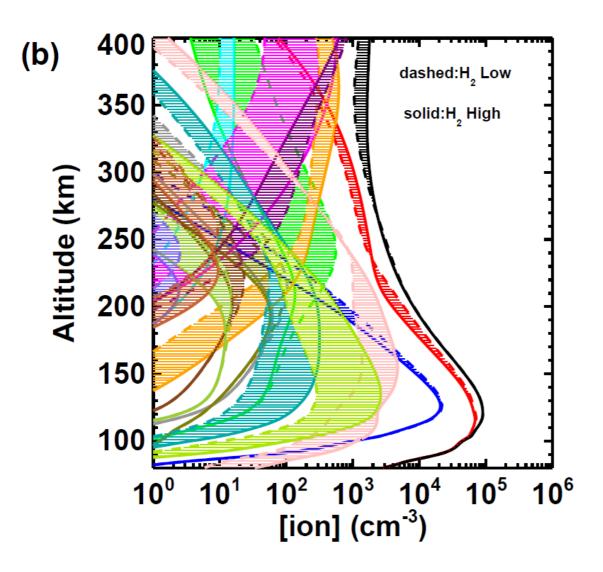
ArH⁺

 N_aH^+

HCO,

HOC*

H⁺



What happens if we add H-species into our model?

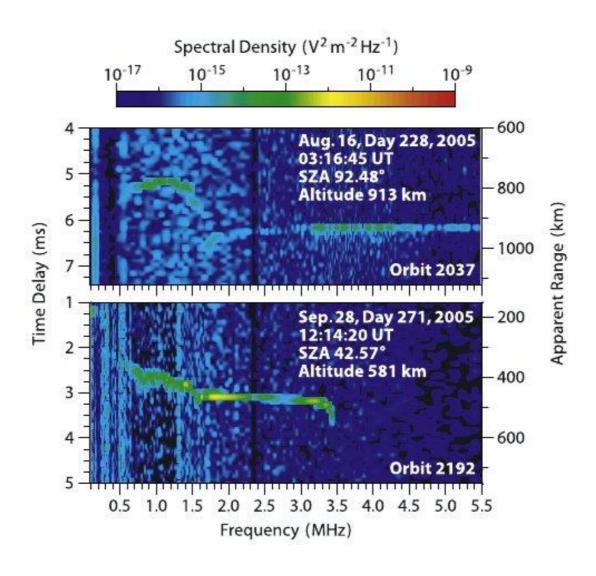
Uncertainties in H2 content and in a key reaction rate $(H^+ + H_2 \rightarrow H_2^+ + H)$

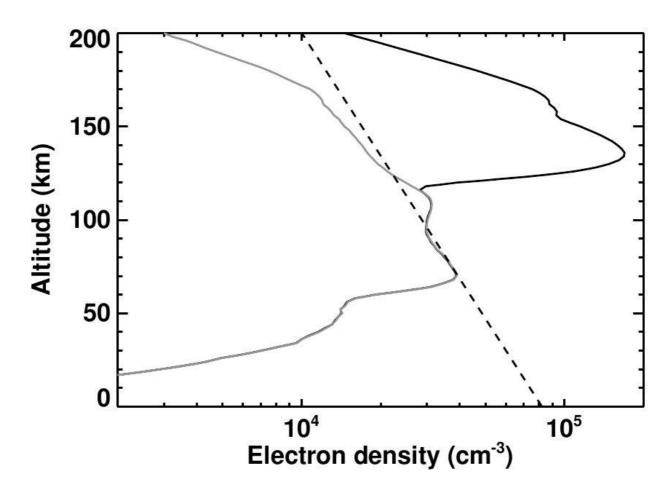
For some regions of parameter space, we find OH+ and related species dominant at high altitudes

Differs from the usual O2+ and O+ dominance

Preliminary!

MARSIS gets blacked out during solar energetic particle (SEP) events





Simulations for an SEP event with and without photoionization Dashed line is simple approximation (I think it does very well!)

Enhanced plasma at 70-100 km Enhanced TEC Massive radio wave absorption for MARSIS

Sheel et al. manuscript submitted to JGR-Space Physics on 15 Dec. (Needed some complicated simulations from Haider's group)